

# Fundamentals of Fluid Dynamics: Waves in Fluids

Introductory Course on Multiphysics Modelling

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(after: D.J. ACHESON's "*Elementary Fluid Dynamics*")

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## 1 Introduction

### 1.1 The notion of wave

#### What is a wave?

A **wave** is the **transport of a disturbance** (or energy, or piece of information) in space not associated with motion of the medium occupying this space as a whole. (Except that electromagnetic waves require no medium !!!)

- The transport is at **finite speed**.

- The shape or form of the **disturbance** is **arbitrary**.
- The disturbance moves with respect to the medium.

Two general classes of wave motion are distinguished:

1. **longitudinal waves** – the disturbance moves parallel to the direction of propagation. *Examples:* sound waves, compressional elastic waves (P-waves in geophysics);
2. **transverse waves** – the disturbance moves perpendicular to the direction of propagation. *Examples:* waves on a string or membrane, shear waves (S-waves in geophysics), water waves, electromagnetic waves.

## 1.2 Basic wave phenomena

**reflection** – change of wave direction from hitting a reflective surface,

**refraction** – change of wave direction from entering a new medium,

**diffraction** – wave circular spreading from entering a small hole (of the wavelength-comparable size), or wave bending around small obstacles,

**interference** – superposition of two waves that come into contact with each other,

**dispersion** – wave splitting up by frequency,

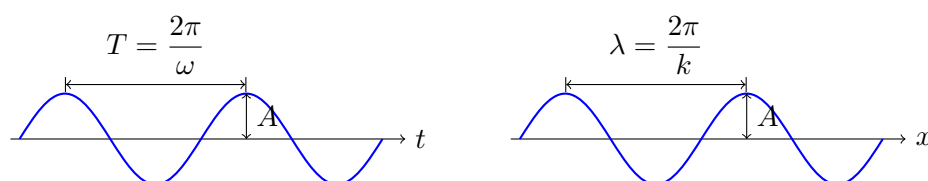
**rectilinear propagation** – the movement of light wave in a straight line.

### Standing wave

A **standing wave**, also known as a **stationary wave**, is a wave that remains in a constant position. This phenomenon can occur:

- when the medium is moving in the opposite direction to the wave,
- (in a stationary medium:) as a result of interference between two waves travelling in opposite directions.

## 1.3 Mathematical description of a traveling wave



**FIGURE 1:** A simple traveling wave in time domain (*left*) and in space (*right*).

### Traveling waves

*Simple wave* or **traveling wave**, sometimes also called *progressive wave*, is a disturbance that varies both with time  $t$  and distance  $x$  in the following way (see Figure 1):

$$\begin{aligned} u(x, t) &= A(x, t) \cos(kx - \omega t + \theta_0) \\ &= A(x, t) \sin\left(kx - \omega t + \underbrace{\theta_0 \pm \frac{\pi}{2}}_{\tilde{\theta}_0}\right) \end{aligned} \quad (1)$$

where  $A$  is the **amplitude**,  $\omega$  and  $k$  denote the **angular frequency** and **wavenumber**, and  $\theta_0$  (or  $\tilde{\theta}_0$ ) is the **initial phase**.

- **Amplitude**  $A$  [e.g. m, Pa, V/m] – a measure of the maximum disturbance in the medium during one wave cycle (the maximum distance from the highest point of the crest to the equilibrium).
- **Phase**  $\theta = kx - \omega t + \theta_0$  [rad], where  $\theta_0$  is the *initial* phase (shift), often ambiguously, called the phase.
- **Period**  $T$  [s] – the time for one complete cycle for an oscillation of a wave.
- **Frequency**  $f$  [Hz] – the number of periods per unit time.

#### Frequency and angular frequency

The **frequency**  $f$  [Hz] represents the number of periods per unit time

$$f = \frac{1}{T}. \quad (2)$$

The **angular frequency**  $\omega$  [Hz] represents the frequency in terms of radians per second. It is related to the frequency by

$$\omega = \frac{2\pi}{T} = 2\pi f. \quad (3)$$

- **Wavelength**  $\lambda$  [m] – the distance between two sequential crests (or troughs).

#### Wavenumber and angular wavenumber

The **wavenumber** is the spatial analogue of frequency, that is, it is the measurement of the number of repeating units of a propagating wave (the number of times a wave has the same phase) per unit of space.

Application of a Fourier transformation on data as a function of time yields a **frequency spectrum**; application on data as a function of position yields a **wavenumber spectrum**.

The **angular wavenumber**  $k$  [ $\frac{1}{\text{m}}$ ], often misleadingly abbreviated as “wave-number”, is defined as

$$k = \frac{2\pi}{\lambda}. \quad (4)$$

There are two velocities that are associated with waves:

1. **Phase velocity** – the rate at which the wave propagates:

$$c = \frac{\omega}{k} = \lambda f. \quad (5)$$

2. **Group velocity** – the velocity at which variations in the shape of the wave’s amplitude (known as the *modulation* or *envelope* of the wave) propagate through space:

$$c_g = \frac{d\omega}{dk}. \quad (6)$$

This is (in most cases) the signal velocity of the waveform, that is, the **rate at which information or energy is transmitted** by the wave. However, if the wave is travelling through an absorptive medium, this does not always hold.

## 2 Water waves

### 2.1 Surface waves on deep water

- Consider **two-dimensional** water waves:  $\mathbf{u} = [u(x, y, t), v(x, y, t), 0]$ .
- Suppose that the flow is **irrotational**:  $\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} = 0$ .
- Therefore, there exists a **velocity potential**  $\phi(x, y, t)$  so that

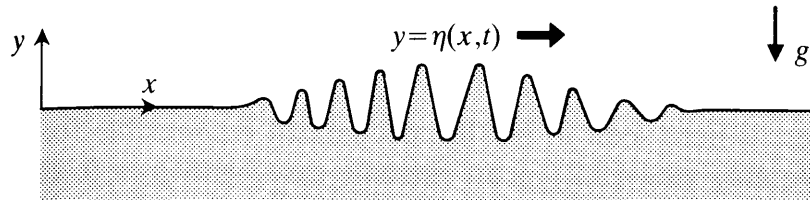
$$u = \frac{\partial \phi}{\partial x}, \quad v = \frac{\partial \phi}{\partial y}. \quad (7)$$

- The fluid is **incompressible**, so by the virtue of the incompressibility condition,  $\nabla \cdot \mathbf{u} = 0$ , the velocity potential  $\phi$  will satisfy **Laplace’s equation**

$$\frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial y^2} = 0. \quad (8)$$

#### Free surface

The fluid motion arises from a deformation of the water surface – which is of major interest (see Figure 2). The equation of this free surface is denoted by  $y = \eta(x, t)$ .



**FIGURE 2:** A deformation on the free surface of water in the form of a wave packet.

### Kinematic condition at the free surface:

Fluid particles on the surface must remain on the surface.

The kinematic condition entails that  $F(x, y, t) = y - \eta(x, t)$  **remains constant** (in fact, zero) for any particular particle on the free surface which means that

$$\frac{DF}{Dt} = \frac{\partial F}{\partial t} + (\mathbf{u} \cdot \nabla)F = 0 \quad \text{on} \quad y = \eta(x, t), \quad (9)$$

and this is equivalent to

$$\frac{\partial \eta}{\partial t} + u \frac{\partial \eta}{\partial x} = v \quad \text{on} \quad y = \eta(x, t). \quad (10)$$

### Pressure condition at the free surface:

The fluid is **inviscid** (by assumption), so the condition at the free surface is simply that the pressure there is equal to the atmospheric pressure  $p_0$ :

$$p = p_0 \quad \text{on} \quad y = \eta(x, t). \quad (11)$$

### Bernoulli's equation for unsteady irrotational flow

If the flow is irrotational (so  $\mathbf{u} = \nabla\phi$  and  $\nabla \times \mathbf{u} = 0$ ), then, by integrating (over the space domain) the **Euler's momentum equation**:

$$\frac{\partial \nabla\phi}{\partial t} = -\nabla \left( \frac{p}{\rho} + \frac{1}{2} \mathbf{u}^2 + \chi \right), \quad (12)$$

the **Bernoulli's equation** is obtained

$$\frac{\partial \phi}{\partial t} + \frac{p}{\rho} + \frac{1}{2} \mathbf{u}^2 + \chi = G(t). \quad (13)$$

Here,  $\chi$  is the gravity potential (in the present context  $\chi = gy$  where  $g$  is the gravity acceleration) and  $G(t)$  is an arbitrary function of time alone (a constant of integration).

Now, by choosing  $G(t)$  in a convenient manner,  $G(t) = \frac{p_0}{\rho}$ , the **pressure condition** may be written as:

$$\frac{\partial \phi}{\partial t} + \frac{1}{2} (u^2 + v^2) + g\eta = 0 \quad \text{on} \quad y = \eta(x, t). \quad (14)$$

### Small-amplitude waves

The free surface displacement  $\eta(x, t)$  and the fluid velocities  $u, v$  are small (in a sense to be made precise later).

#### Linearization of the kinematic condition

$$v = \frac{\partial \eta}{\partial t} + \underbrace{u \frac{\partial \eta}{\partial x}}_{\text{small}} \rightarrow v(x, \eta, t) = \frac{\partial \eta}{\partial t}$$

$$\xrightarrow{\text{Taylor series}} v(x, 0, t) + \underbrace{\eta \frac{\partial v}{\partial y}(x, 0, t) + \dots}_{\text{small}} = \frac{\partial \eta}{\partial t} \quad (15)$$

$$\rightarrow v(x, 0, t) = \frac{\partial \eta}{\partial t} \quad \xrightarrow{v = \frac{\partial \phi}{\partial y}} \left( \frac{\partial \phi}{\partial y} = \frac{\partial \eta}{\partial t} \text{ on } y = 0. \right)$$

#### Linearization of the pressure condition

$$\frac{\partial \phi}{\partial t} + \underbrace{\frac{1}{2}(u^2 + v^2)}_{\text{small}} + g\eta = 0 \rightarrow \left( \frac{\partial \phi}{\partial t} + g\eta = 0 \text{ on } y = 0. \right) \quad (16)$$

### A sinusoidal travelling wave solution

The **free surface** is of the form

$$\eta = A \cos(kx - \omega t), \quad (17)$$

where  $A$  is the **amplitude** of the surface displacement,  $\omega$  is the **circular frequency**, and  $k$  is the **circular wavenumber**.

#### The corresponding velocity potential is

$$\phi = q(y) \sin(kx - \omega t). \quad (18)$$

It satisfies the Laplace's equation,  $\frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial y^2} = 0$ .

Therefore,  $q(y)$  must satisfy  $q'' - k^2 q = 0$ , the general solution of which is

$$q = C \exp(ky) + D \exp(-ky). \quad (19)$$

For *deep* water waves  $D = 0$  (if  $k > 0$  which may be assumed without loss of generality) in order that the velocity be bounded as  $y \rightarrow -\infty$ . Therefore, the velocity potential for *deep* water waves is

$$\phi = C \exp(ky) \sin(kx - \omega t). \quad (20)$$

■ Now, the (linearized) **free surface conditions** yield what follows:

1. the **kinematic condition** ( $\frac{\partial \phi}{\partial y} = \frac{\partial \eta}{\partial t}$  on  $y = 0$ ):

$$C k = A \omega \quad \rightarrow \quad \boxed{\phi = \frac{A \omega}{k} \exp(k y) \sin(k x - \omega t)}, \quad (21)$$

2. the **pressure condition** ( $\frac{\partial \phi}{\partial t} + g \eta = 0$  on  $y = 0$ ):

$$-C \omega + g A = 0 \quad \rightarrow \quad \boxed{\omega^2 = g k}. \quad (\text{dispersion relation!}) \quad (22)$$

The **fluid velocity** components:

$$u = A \omega \exp(k y) \cos(k x - \omega t), \quad v = A \omega \exp(k y) \sin(k x - \omega t). \quad (23)$$

### Particle paths

Any particle departs only a **small amount**  $(X, Y)$  **from its mean position**  $(x, y)$ . Therefore, its position as a function of time may be found by integrating  $u = \frac{dX}{dt}$  and  $v = \frac{dY}{dt}$ ; whence:

$$X(t) = -A \exp(k y) \sin(k x - \omega t), \quad Y(t) = A \exp(k y) \cos(k x - \omega t). \quad (24)$$

Figure 3 presents particle paths for a wave on deep water. One may observe what follows:

- Particle **paths are circular**.
- The **radius** of the path circles,  $A \exp(k y)$ , **decrease exponentially with depth**. So do the **fluid velocities**.
- Virtually all the **energy** of a surface water wave is **contained within half a wavelength below the surface**.

### Effects of finite depth

If the **fluid is bonded below** by a rigid plane  $y = -h$ , so that

$$v = \frac{\partial \phi}{\partial y} = 0 \quad \text{at } y = -h, \quad (25)$$

the **dispersion relation** and the **phase speed** are as follows:

$$\omega^2 = g k \tanh(k h), \quad c^2 = \frac{g}{k} \tanh(k h). \quad (26)$$

Figure 4 shows the phase speed of waves in water of uniform depth  $h$  in function of the wavelength  $\lambda = \frac{2\pi}{k}$ . There are two limit cases:

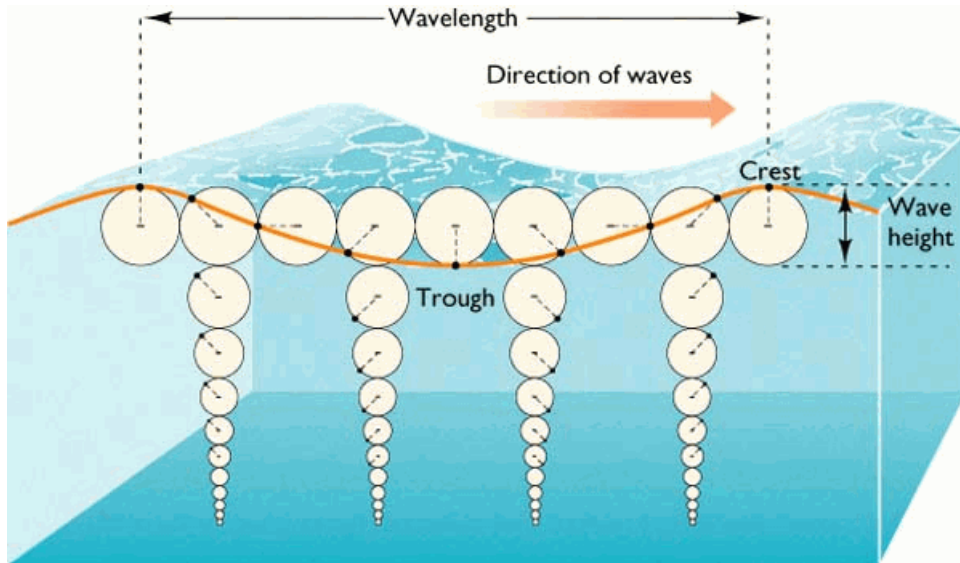


FIGURE 3: Deep water wave and circular particle paths (Wikipedia).

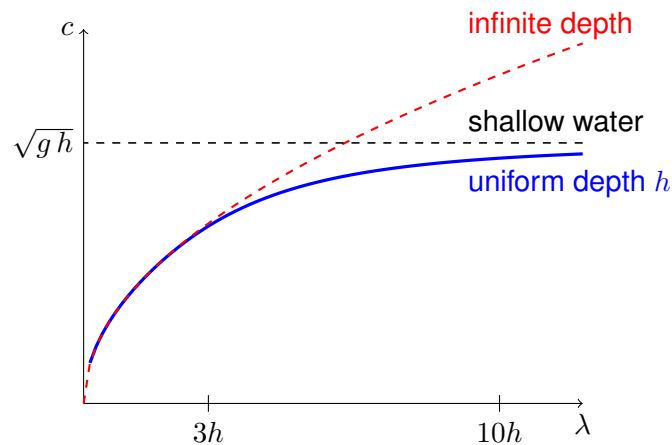


FIGURE 4: The phase speed of gravity waves in water of uniform depth  $h$ .

1.  $h \gg \lambda$  (**infinite depth**):  $kh = 2\pi \frac{h}{\lambda}$  is large and  $\tanh(kh) \approx 1$ , so  $c^2 = \frac{g}{k}$ . In practice, this is a good approximation if  $h > \frac{1}{3}\lambda$ .
2.  $h \ll \lambda/2\pi$  (**shallow water**):  $kh \ll 1$  and  $\tanh(kh) \approx kh$ , so  $c^2 = gh$ , which means that  $c$  is independent of  $k$  in this limit. Thus, the **gravity waves in shallow water are non-dispersive**.

## 2.2 Dispersion and the group velocity

### Dispersion of waves

Dispersion of waves is the phenomenon that the **phase velocity of a wave depends on its frequency**.



There are generally two sources of dispersion:

1. the **material dispersion** comes from a frequency-dependent response of a material to waves
2. the **waveguide dispersion** occurs when the speed of a wave in a waveguide depends on its frequency for geometric reasons, independent of any frequency-dependence of the materials from which it is constructed.

### Dispersion relation

The dispersion exists when the (angular) frequency is related to the wavenumber in a non-linear way:

$$\boxed{\omega = \omega(k)} = c(k) k, \quad c = c(k) = \frac{\omega(k)}{k}. \quad (27)$$

If  $\omega(k)$  is a **linear** function of  $k$  then  $c$  is **constant** and the medium is **non-dispersive**.

Dispersion relations for waves on water surface:

- ▶ deep water waves:  $\omega = \sqrt{g k}$ ,  $c = \sqrt{\frac{g}{k}}$ .
- ▶ finite depth waves:  $\omega = \sqrt{g k \tanh(k h)}$ ,  $c = \sqrt{\frac{g}{k} \tanh(k h)}$ .
- ▶ shallow water waves:  $\omega = \sqrt{g h} k$ ,  $c = \sqrt{g h} \rightarrow$  non-dispersive!

### Group and phase velocity

Two fundamental velocities of wave propagation, namely, the group velocity  $c_g$  and the phase velocity  $c$ , are defined as follows:

$$\boxed{c_g = \frac{d\omega}{dk}}, \quad c = \frac{\omega}{k}. \quad (28)$$

- In **dispersive systems** both velocities are different and frequency-dependent (i.e., wavenumber-dependent):  $c_g = c_g(k)$  and  $c = c(k)$ .
- In **non-dispersive systems** they are equal and constant:  $c_g = c$ .

**Important properties of the group velocity:**

1. At this velocity the isolated **wave packet** travels as *a whole*.

**Discussion for a wave packet** (see Figure 5): for  $k$  in the neighbourhood of  $k_0$

$$\omega(k) \approx \omega(k_0) + (k - k_0) c_g, \quad \text{where } c_g = \left. \frac{d\omega}{dk} \right|_{k=k_0}, \quad (29)$$

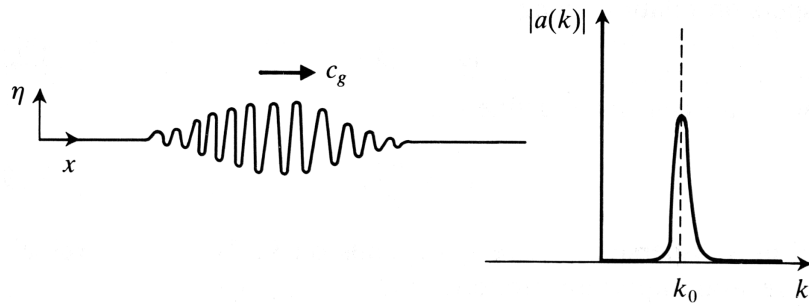


FIGURE 5: A wave packet and its spectrum.

and  $\omega(k) = 0$  outside the neighbourhood; the Fourier integral equals

$$\eta(x, t) = \text{Re} \left[ \int_{-\infty}^{\infty} a(k) \exp(i(kx - \omega t)) dk \right] \quad \leftarrow \text{(for a general disturbance)}$$

$$\approx \text{Re} \left[ \underbrace{\exp(i(k_0 x - \omega(k_0) t))}_{\text{a pure harmonic wave}} \int_{-\infty}^{\infty} \underbrace{a(k) \exp(i(k - k_0)(x - c_g t))}_{\text{a function of } (x - c_g t)} dk \right]. \quad (30)$$

2. The **energy is transported** at the group velocity (by waves of a given wavelength).
3. One must travel at the group velocity to see the waves of the same wavelength.

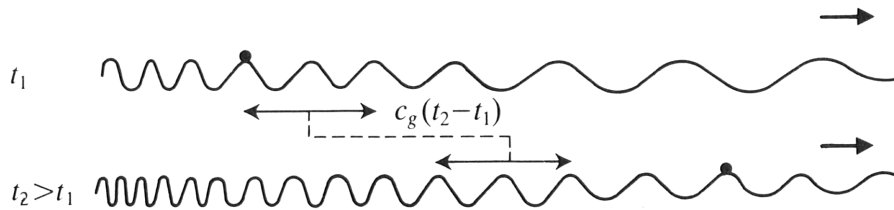


FIGURE 6: A train of waves.

A slowly varying **wavetrain** (see Figure 6) can be written as

$$\eta(x, t) = \text{Re} [A(x, t) \exp(i\theta(x, t))], \quad (31)$$

where the **phase function**  $\theta(x, t)$  describes the oscillatory aspect of the wave, while  $A(x, t)$  describes the gradual modulation of its amplitude.

The *local* wavenumber and frequency are defined by

$$k = \frac{\partial \theta}{\partial x}, \quad \omega = -\frac{\partial \theta}{\partial t}. \quad (32)$$

For purely sinusoidal wave  $\theta = kx - \omega t$ , where  $k$  and  $\omega$  are constants. In general,  $k$  and  $\omega$  are functions of  $x$  and  $t$ . It follows immediately that

$$\frac{\partial k}{\partial t} + \frac{\partial \omega}{\partial x} = 0 \quad \longrightarrow \quad \frac{\partial k}{\partial t} + \frac{d\omega}{dk} \frac{\partial k}{\partial x} = \frac{\partial k}{\partial t} + c_g(k) \frac{\partial k}{\partial x} = 0 \quad (33)$$

which means that  $k(x, t)$  is constant for an observer moving with the velocity  $c_g(k)$ .

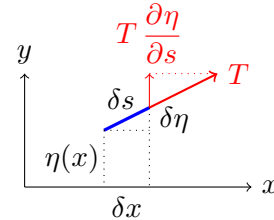
## 2.3 Capillary waves

### Surface tension

A **surface tension force**  $T$  [ $\frac{\text{N}}{\text{m}}$ ] is a force per unit length, directed tangentially to the surface, acting on a line drawn parallel to the wavecrests.

- The **vertical component** of surface tension force equals  $T \frac{\partial \eta}{\partial s}$ , where  $s$  denotes the distance along the surface.

- For **small wave amplitudes**  $\delta s \approx \delta x$ , and then  $T \frac{\partial \eta}{\partial s} \approx T \frac{\partial \eta}{\partial x}$ .



- A small portion of surface of length  $\delta x$  will experience surface tension at both ends, so the **net upward force** on it will be

$$T \frac{\partial \eta}{\partial x} \Big|_{x+\delta x} - T \frac{\partial \eta}{\partial x} \Big|_x = T \frac{\partial^2 \eta}{\partial x^2} \delta x \quad (34)$$

- Therefore, an upward force **per unit area of surface** is  $T \frac{\partial^2 \eta}{\partial x^2}$ .

### Local equilibrium at the free surface

The net upward force per unit area of surface,  $T \frac{\partial^2 \eta}{\partial x^2}$ , must be balanced by the difference between the atmospheric pressure  $p_0$  and the pressure  $p$  in the fluid just below the surface:

$$p_0 - p = T \frac{\partial^2 \eta}{\partial x^2} \quad \text{on} \quad y = \eta(x, t). \quad (35)$$

This **pressure condition** at the free surface takes into consideration the **effects of surface tension**. The kinematic condition remains the same: fluid particles cannot leave the surface.

### Linearized free surface conditions (with surface tension effects)

For small amplitude waves:

$$\frac{\partial \phi}{\partial y} = \frac{\partial \eta}{\partial t}, \quad \frac{\partial \phi}{\partial t} + g\eta = \frac{T}{\rho} \frac{\partial^2 \eta}{\partial x^2} \quad \text{on} \quad y = 0. \quad (36)$$

Notice that the right-hand-side term of the pressure condition results from a **surface tension**.

A sinusoidal travelling wave solution  $\eta = A \cos(kx - \omega t)$  leads now to a new **dispersion relation**

$$\omega^2 = gk + \frac{Tk^3}{\rho}. \quad (37)$$

As a consequence, the **phase** and **group velocities** include now the surface tension effect:

$$c = \frac{\omega}{k} = \sqrt{\frac{g}{k} + \frac{Tk}{\rho}}, \quad c_g = \frac{d\omega}{dk} = \frac{g + 3Tk^2/\rho}{2\sqrt{gk + Tk^3/\rho}}. \quad (38)$$

### Surface tension importance parameter

The relative importance of surface tension and gravitational forces in a fluid is measured by the following parameter

$$\beta = \frac{Tk^2}{\rho g}. \quad (39)$$

(The so-called *Bond number*  $= \frac{\rho g L^2}{T}$ ; it equals  $\frac{4\pi^2}{\beta}$  if  $L = \lambda$ .)

Now, the dispersion relation, as well as the phase and group velocities can be written as

$$\omega^2 = gk(1 + \beta), \quad c = \sqrt{\frac{g}{k}(1 + \beta)}, \quad c_g = \frac{g(1 + 3\beta)}{2\sqrt{gk(1 + \beta)}}. \quad (40)$$

Depending on the parameter  $\beta$ , two extreme cases are distinguished:

1.  $\beta \ll 1$ : the effects of surface tension are negligible – the waves are **gravity waves** for which

$$\omega^2 = gk, \quad c = \sqrt{\frac{g}{k}} = \sqrt{\frac{g\lambda}{2\pi}}, \quad c_g = \frac{c}{2}. \quad (41)$$

2.  $\beta \gg 1$ : the waves are essentially **capillary waves** for which

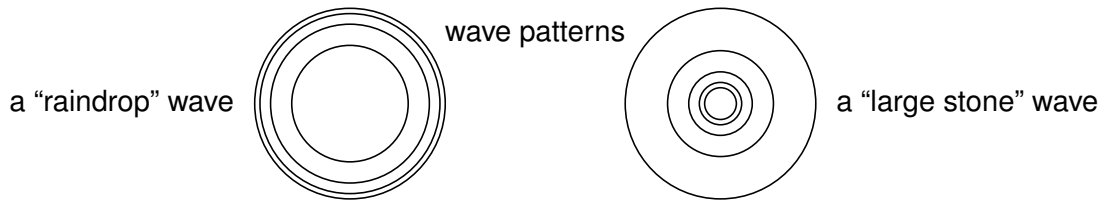
$$\omega^2 = gk\beta = \frac{Tk^3}{\rho}, \quad c = \sqrt{\frac{g}{k}\beta} = \sqrt{\frac{Tk}{\rho}} = \sqrt{\frac{2\pi T}{\rho\lambda}}, \quad c_g = \frac{g3\beta}{2\sqrt{gk\beta}} = \frac{3}{2}c. \quad (42)$$

### CAPILLARY WAVES:

- **short waves travel faster**,
- the group velocity exceeds the phase velocity,  $c_g > c$ ,
- the **wavecrests move backward** through a wave packet as it moves along as a whole.

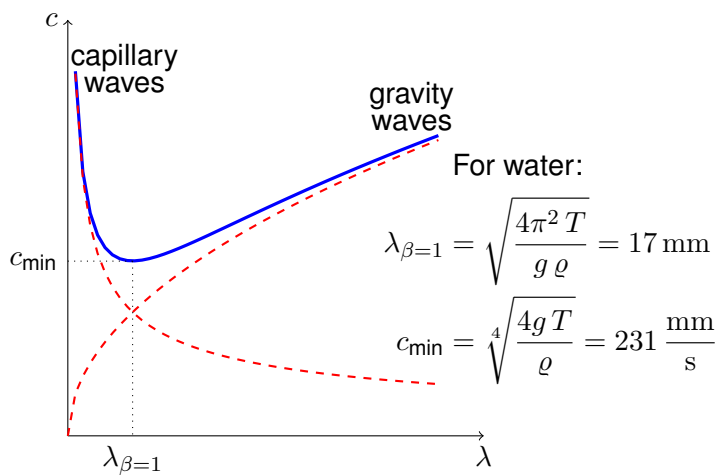
### GRAVITY WAVES:

- **long waves travel faster**,
- the group velocity is less than the phase velocity,  $c_g < c$ ,
- the **wavecrests move faster** than a wave packet.



The **capillary effects** predominate when **raindrops** fall on a pond, and as short waves travel faster the **wavelength decreases with radius** at any particular time.

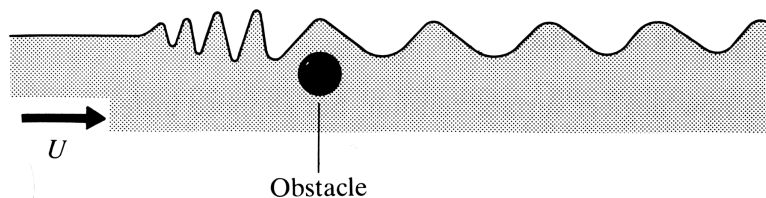
The **effects of gravity** predominate when a **large stone** is dropped into a pond (on account of the longer wavelengths involved), and as long waves travel faster the **wavelength increases with radius** at any particular time.



**FIGURE 7:** The phase speed for capillary-gravity waves.

For  $\beta \approx 1$  both effects (the surface tension and gravity) are significant and the waves are **capillary-gravity waves**. Figure 7 presents the phase speed of such waves depending on the wavelength  $\lambda = \frac{2\pi}{k}$ . Notice that the speed reaches its minimum,  $c = c_{\min}$ , for such  $\lambda$  that  $\beta = 1$ . For water at 20°C (when  $T = 7.29 \times 10^{-4} \frac{\text{N}}{\text{m}}$  and  $\rho = 998 \frac{\text{kg}}{\text{m}^3}$ ) this is when the wavelength is about 17 mm.

**Example: Uniform flow past a submerged obstacle**



**FIGURE 8:** Stationary waves generated by uniform flow, speed  $U$ , past a submerged obstacle.

Figure 8 presents stationary waves generated by uniform flow past a submerged obstacle. Two cases are distinguished with respect to the flow speed  $U$ , namely:

1.  $U < c_{\min}$  – there are no steady waves generated by the obstacle;
2.  $U > c_{\min}$  – there are **two values** of  $\lambda$  ( $\lambda_1 > \lambda_2$ ) for which  $c = U$ :
  - $\lambda_1$  – **the larger value** represents a **gravity wave**:
    - the corresponding group velocity is less than  $c$ ,
    - the **energy** of this relatively long-wavelength disturbance **is carried downstream** of the obstacle.
  - $\lambda_2$  – **the smaller value** represents a **capillary wave**:
    - the corresponding group velocity is greater than  $c$ ,
    - the **energy** of this relatively short-wavelength disturbance **is carried upstream** of the obstacle, where it is rather quickly dissipated by viscous effects, on account of the short wavelength (in fact, each wave-crest is at rest, but relative to still water it is travelling upstream with speed  $U$ ).

## 2.4 Shallow-water finite-amplitude waves

### Assumptions:

- The **amplitudes of waves are finite**, that is, *not* (infinitesimally) small compared with the depth; therefore, the **linearized theory does not apply**.
- A typical value  $h_0$  of depth  $h(x, t)$  is much smaller than a typical horizontal length scale  $L$  of the wave (see Figure 9), that is:  $h_0 \ll L$ . This is the basis of the so-called **shallow-water approximation**.

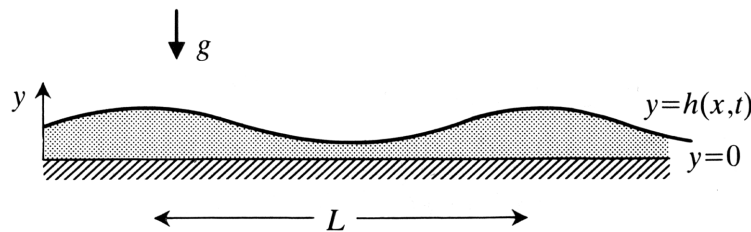


FIGURE 9: Finite-amplitude wave on shallow water.

- ▶ The full (nonlinear) 2-D equations are:

$$\frac{Du}{Dt} = -\frac{1}{\rho} \frac{\partial p}{\partial x}, \quad \frac{Dv}{Dt} = -\frac{1}{\rho} \frac{\partial p}{\partial y} - g, \quad \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0. \quad (43)$$

- ▶ In the shallow-water approximation (when  $h_0 \ll L$ ) the vertical component of acceleration can be neglected in comparison with the gravitational acceleration:

$$\frac{Dv}{Dt} \ll g \quad \rightarrow \quad 0 = -\frac{1}{\rho} \frac{\partial p}{\partial y} - g \quad \rightarrow \quad \frac{\partial p}{\partial y} = \rho g. \quad (44)$$

Integrating and applying the condition  $p = p_0$  at  $y = h(x, t)$  gives

$$p(x, y, t) = p_0 - \rho g [y - h(x, t)]. \quad (45)$$

This is used for the equation for the horizontal component of acceleration:

$$\frac{Du}{Dt} = -g \frac{\partial h}{\partial x} \xrightarrow{\frac{\partial u}{\partial y} = 0} \boxed{\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} = -g \frac{\partial h}{\partial x}} \quad (46)$$

where  $u = u(x, t)$  and  $h = h(x, t)$ .

► A second equation linking  $u$  and  $h$  may be obtained as follows:

$$\frac{\partial v}{\partial y} = -\frac{\partial u}{\partial x} \rightarrow v(x, y, t) = -\frac{\partial u(x, t)}{\partial x} y + f(x, t) \xrightarrow{v=0 \text{ at } y=0} v = -\frac{\partial u}{\partial x} y, \quad (47)$$

and using the **kinematic condition at the free surface** – fluid particles on the surface must remain on it, so the vertical component of velocity  $v$  equals the rate of change of the depth  $h$  when moving with the horizontal velocity  $u$ :

$$v = \frac{\partial h}{\partial t} + u \frac{\partial h}{\partial x} \quad \text{at } y = h(x, t) \rightarrow \boxed{\frac{\partial h}{\partial t} + u \frac{\partial h}{\partial x} + h \frac{\partial u}{\partial x} = 0}. \quad (48)$$

### Shallow-water equations

Nonlinear equations for the horizontal component of velocity  $u = u(x, t)$  and the depth  $h = h(x, t)$  of finite-amplitude waves on shallow water:

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + g \frac{\partial h}{\partial x} = 0, \quad \frac{\partial h}{\partial t} + u \frac{\partial h}{\partial x} + h \frac{\partial u}{\partial x} = 0. \quad (49)$$

(The vertical component of velocity is  $v(x, y, t) = -\frac{\partial u}{\partial x} y$ .)

On introducing the new variable  $c(x, t) = \sqrt{gh}$  and then adding and subtracting the two equations the form suited to treatment by the *method of characteristics* is obtained

$$\left[ \frac{\partial}{\partial t} + (u + c) \frac{\partial}{\partial x} \right] (u + 2c) = 0, \quad \left[ \frac{\partial}{\partial t} + (u - c) \frac{\partial}{\partial x} \right] (u - 2c) = 0. \quad (50)$$

Let  $x = x(s)$ ,  $t = t(s)$  be a **characteristic curve** defined parametrically ( $s$  is the parameter) in the  $x$ - $t$  plane and starting at some point  $(x_0, t_0)$ . In fact, two such (families of) characteristic curves are defined such that:

$$\frac{dt}{ds} = 1, \quad \frac{dx}{ds} = u \pm c. \quad (51)$$

This (with  $+$ ) is used for the first and (with  $-$ ) for the second equation:

$$\left[ \frac{dt}{ds} \frac{\partial}{\partial t} + \frac{dx}{ds} \frac{\partial}{\partial x} \right] (u \pm 2c) = 0 \xrightarrow{\text{the chain rule}} \boxed{\frac{d}{ds} (u \pm 2c) = 0}. \quad (52)$$

**General property:**  $u \pm 2c$  is constant along 'positive'/'negative' characteristic curves defined by  $\frac{dx}{dt} = u \pm c$ .

Within the framework of the theory of finite-amplitude waves on shallow water the following problems can be solved:

- the dam-break flow,
- the formation of a bore,
- the hydraulic jump.

## 3 Sound waves

### 3.1 Introduction

**Sound waves** propagate due to the **compressibility** of a medium ( $\nabla \cdot \mathbf{u} \neq 0$ ). Depending on frequency one can distinguish:

- **infrasound waves** – below 20 Hz,
- **acoustic waves** – from 20 Hz to 20 kHz,
- **ultrasound waves** – above 20 kHz.

**Acoustics** deals with vibrations and waves in compressible continua in the **audible frequency range**, that is, from 20 Hz (16 Hz) to 20 000 Hz.

Types of waves in compressible continua:

- an **inviscid compressible fluid** – (only) longitudinal waves,
- an infinite **isotropic solid** – longitudinal and shear waves,
- an **anisotropic solid** – wave propagation is more complex.

### 3.2 Acoustic wave equation

**Assumptions:**

- Gravitational forces can be neglected so that the equilibrium (undisturbed-state) pressure and density take on uniform values,  $p_0$  and  $\rho_0$ , throughout the fluid.
- Dissipative effects, that is viscosity and heat conduction, are neglected.
- The medium (fluid) is homogeneous, isotropic, and perfectly elastic.



**Small-amplitudes assumption**

Particle velocity is small, and there are only very small perturbations (fluctuations) to the equilibrium pressure and density:

$$\mathbf{u} - \text{small}, \quad p = p_0 + \tilde{p} \quad (\tilde{p} - \text{small}), \quad \varrho = \varrho_0 + \tilde{\varrho} \quad (\tilde{\varrho} - \text{small}). \quad (53)$$

The pressure fluctuations field  $\tilde{p}$  is called the **acoustic pressure**.

**Momentum equation** (Euler's equation):

$$\varrho \left( \frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = -\nabla p \quad \xrightarrow{\text{linearization}} \quad \varrho_0 \frac{\partial \mathbf{u}}{\partial t} = -\nabla p. \quad (54)$$

Notice that  $\nabla p = \nabla(p_0 + \tilde{p}) = \nabla \tilde{p}$ .

**Continuity equation:**

$$\frac{\partial \varrho}{\partial t} + \nabla \cdot (\varrho \mathbf{u}) = 0 \quad \xrightarrow{\text{linearization}} \quad \frac{\partial \tilde{\varrho}}{\partial t} + \varrho_0 \nabla \cdot \mathbf{u} = 0. \quad (55)$$

Using divergence operation for the linearized momentum equation and time-differentiation for the linearized continuity equation yields:

$$\frac{\partial^2 \tilde{\varrho}}{\partial t^2} - \Delta p = 0. \quad (56)$$

**Constitutive relation:**

$$p = p(\tilde{\varrho}) \quad \rightarrow \quad \frac{\partial p}{\partial t} = \frac{\partial p}{\partial \tilde{\varrho}} \frac{\partial \tilde{\varrho}}{\partial t} \quad \rightarrow \quad \frac{\partial^2 p}{\partial t^2} = \frac{1}{c_0^2} \frac{\partial^2 p}{\partial t^2} \quad \text{where } c_0^2 = \frac{\partial p}{\partial \tilde{\varrho}}. \quad (57)$$

**Wave equation for the pressure field**

$$\left( \frac{1}{c_0^2} \frac{\partial^2 p}{\partial t^2} - \Delta p = 0 \right) \quad \text{where } c_0 = \sqrt{\frac{\partial p}{\partial \tilde{\varrho}}} \quad (58)$$

is the **acoustic wave velocity** (or the **speed of sound**). Notice that the acoustic pressure  $\tilde{p}$  can be used here instead of  $p$ . Moreover, the wave equation for the density-fluctuation field  $\tilde{\varrho}$  (or for the compression field  $\tilde{\varrho}/\varrho_0$ ), for the velocity potential  $\phi$ , and for the velocity field  $\mathbf{u}$  can be derived analogously.

**3.3 The speed of sound**

**Inviscid isotropic elastic liquid.** The pressure in an inviscid liquid depends on the volume dilatation  $\text{tr } \varepsilon$ :

$$p = -K \text{tr } \varepsilon, \quad (59)$$

where  $K$  is the bulk modulus. Now,

$$\frac{\partial p}{\partial t} = -K \operatorname{tr} \frac{\partial \boldsymbol{\varepsilon}}{\partial t} = -K \nabla \cdot \mathbf{u} \quad \begin{array}{l} \nabla \cdot \mathbf{u} = -\frac{1}{\rho_0} \frac{\partial \tilde{\rho}}{\partial t} \\ \text{Lin. Cont. Eq.} \end{array} \rightarrow \frac{\partial p}{\partial t} = \frac{K}{\rho_0} \frac{\partial \tilde{\rho}}{\partial t} \quad (60)$$

which means that the speed of sound  $c_0 = \sqrt{\partial p / \partial \tilde{\rho}}$  is given by the well-known formula:

$$c_0 = \sqrt{\frac{K}{\rho_0}}. \quad (61)$$

**Perfect gas.** The determination of speed of sound in a perfect gas is complicated and requires the use of thermodynamic considerations. The final result is

$$c_0 = \sqrt{\gamma \frac{p_0}{\rho_0}} = \sqrt{\gamma R T_0}, \quad (62)$$

where  $\gamma$  denotes the ratio of specific heats ( $\gamma = 1.4$  for air),  $R$  is the universal gas constant, and  $T_0$  is the (isothermal) temperature.

► For air at 20°C and normal atmospheric pressure:  $c_0 = 343 \frac{\text{m}}{\text{s}}$ .

### 3.4 Sub- and supersonic flow

A steady, unseparated, **compressible flow** past a thin airfoil may be written in the form

$$u = U + \frac{\partial \phi}{\partial x}, \quad v = \frac{\partial \phi}{\partial y}, \quad (63)$$

where the **velocity potential**  $\phi$  for the small disturbance to the uniform flow  $U$  satisfies

$$(1 - M^2) \frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial y^2} = 0, \quad \text{where} \quad M = \frac{U}{c_0} \quad (64)$$

is the **Mach number** defined as the ratio of the speed of free stream to the speed of sound.

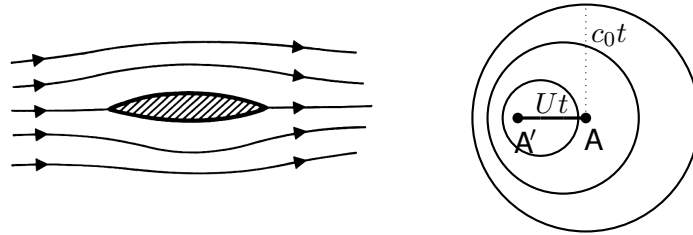
► If  $M^2 \ll 1$  that gives the Laplace equation which is the result that arises for **incompressible theory** (i.e., using  $\nabla \cdot \mathbf{u} = 0$ ).

► Otherwise, three cases can be distinguished:

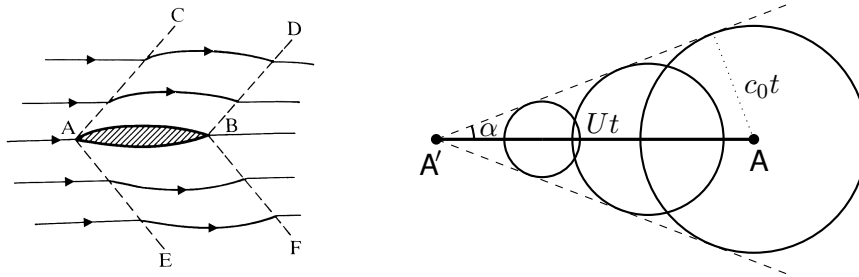
1.  $M < 1$  – the **subsonic flow** (see Figure 10):

- there is **some disturbance to the oncoming flow** at all distances from the wing (even though it is very small when the distance is large);
- the **drag is zero** (inviscid theory) and the lift =  $\frac{\text{lift}_{\text{incompressible}}}{\sqrt{1-M^2}}$ .

2.  $M > 1$  – the **supersonic flow** (see Figure 11):



**FIGURE 10:** (Left:) Subsonic flow past a thin wing at zero incidence. (Right:) Acoustic radiation by a body moving subsonically ( $M = 0.6$ ).



**FIGURE 11:** (Left:) Supersonic flow past a thin wing at zero incidence. (Right:) Acoustic radiation by a body moving supersonically ( $M = 2.8$ ).

- there is **no disturbance to the oncoming stream** except between the **Mach lines** extending from the ends of the airfoil and making the angle  $\alpha = \arcsin\left(\frac{1}{M}\right)$  with the uniform stream;
- the **drag is not zero** – it arises because of the sound wave energy which the wing radiates to infinity between the Mach lines.

### 3. $M \approx 1$ – the **sound barrier**:

- sub- and supersonic theory is not valid;
- nonetheless, it indicates that the wing is subject to a destructive effect of exceptionally **large aerodynamic forces**.